



Max-Planck-Institut  
für Radioastronomie

# Back-ends for THz heterodyne systems:

## Fast Fourier Transform Spectrometer (FFTS)

Bernd Klein<sup>1,2</sup>

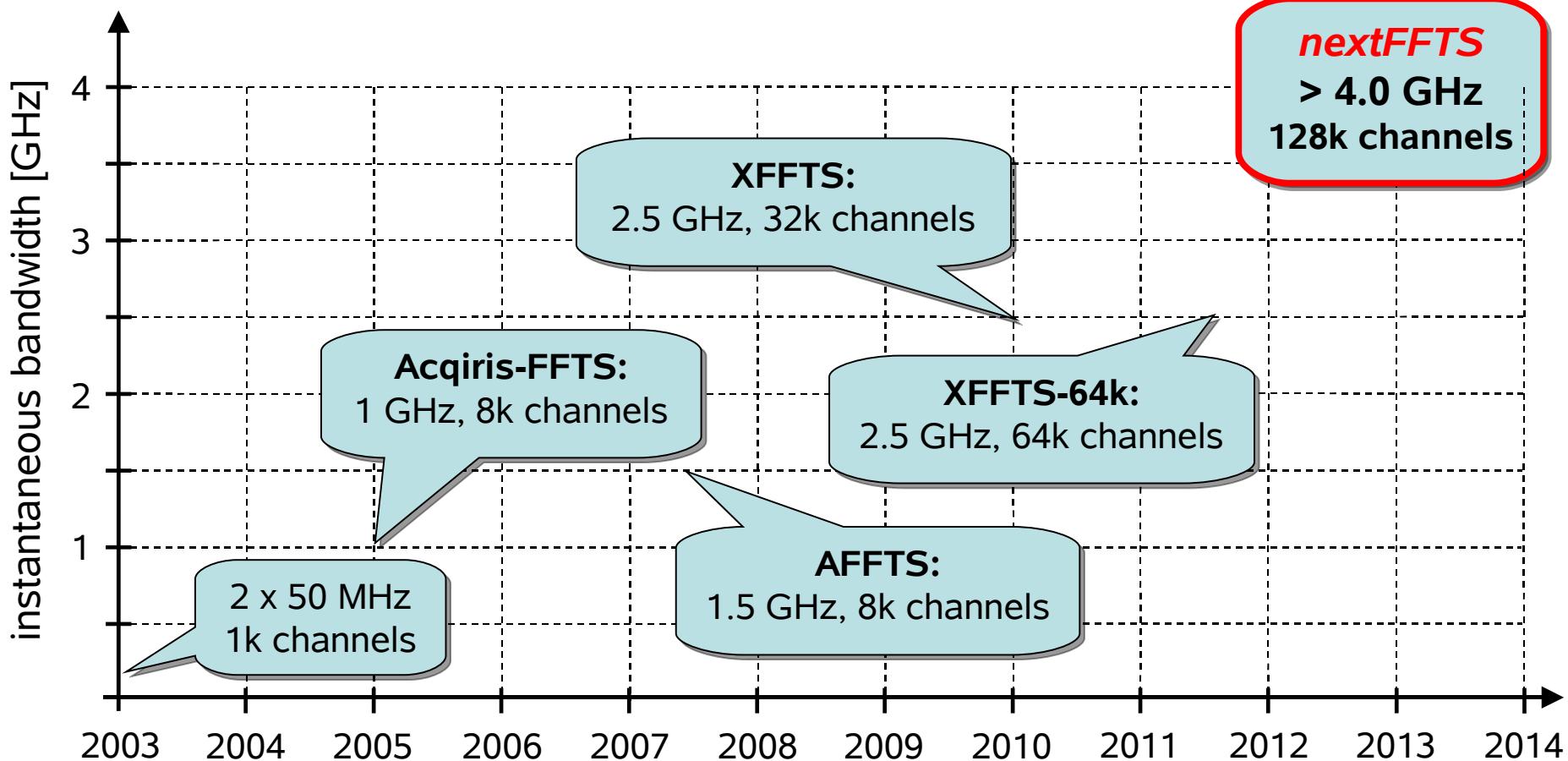
<sup>1</sup>*Max-Planck-Institut für Radioastronomie, Bonn, Germany*

<sup>2</sup>*University of Applied Science Bonn-Rhein-Sieg, Germany*

2014-01-20



## FFTS development history





# Fast Fourier Transform Spectrometer (FFTS)

1. Generation: AFFTS – 1.5 GHz BW, 8K channels

2. Generation: XFFTS – 2.5 GHz BW, 32K channels  
XFFTS2 – 2.5 GHz BW, 64K channels

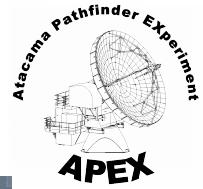
3. Generation: *nextFFTS* – 4 GHz BW, 128K channels  
+ IF sampling (4 – 8 GHz)  
+ digital sideband separation



Max-Planck-Institut  
für Radioastronomie

## APEX :: AFFTS & XFFTS

DIGITALLABOR  
01001010





Max-Planck-Institut  
für Radioastronomie

## APEX :: Flash<sup>+</sup> + XIF + XFFTS / XFFTS2

DIGITALLABOR  
0100110 01000110 01010100 01010011 – 01000010 01001011



- ▶ **FLASH<sup>+</sup> 345:** 4 x XFFTS2 → 4 x 2.5 GHz BW, 4 x 64K channels
- ▶ **FLASH<sup>+</sup> 460:** 4 x XFFTS → 4 x 2.5 GHz BW, 4 x 32K channels





Max-Planck-Institut  
für Radioastronomie

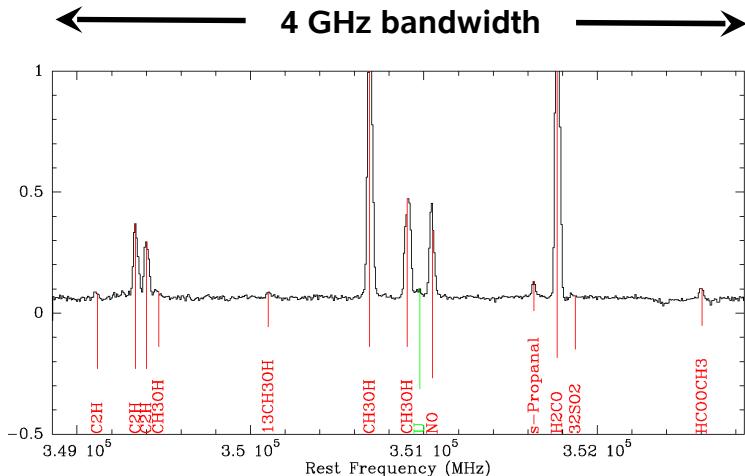
## APEX :: Flash<sup>+</sup>345 + XIF + XFFTS2

DIGITAL LABOR  
01001010

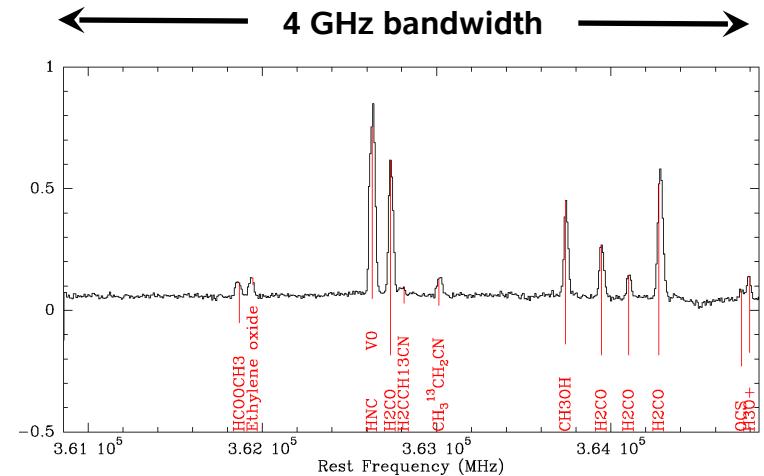


► **8 GHz of bandwidth and ~210'000 spectral channels** ◀

*lower sideband*



*upper sideband*



← 12 GHz spacing →

**Flash<sup>+</sup>345:** updated receiver with IRAM 2SB SIS mixer





Max-Planck-Institut  
für Radioastronomie

## SOFIA / GREAT :: XFFTS

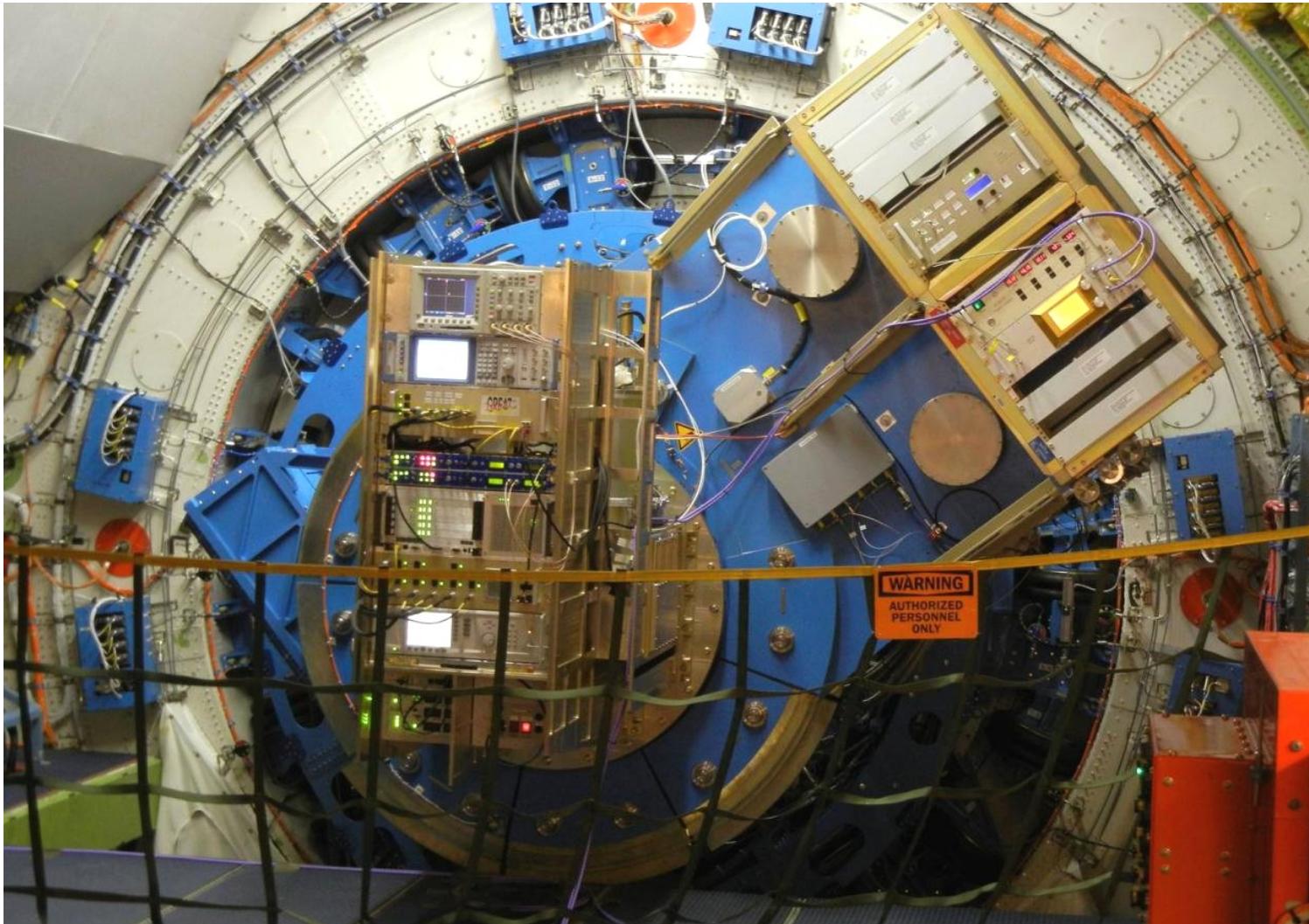
DIGITALLABOR  
0100110 01000110 01010100 01010011 – 01000010 01001011





## SOFIA / GREAT :: XFFTS

Max-Planck-Institut  
für Radioastronomie



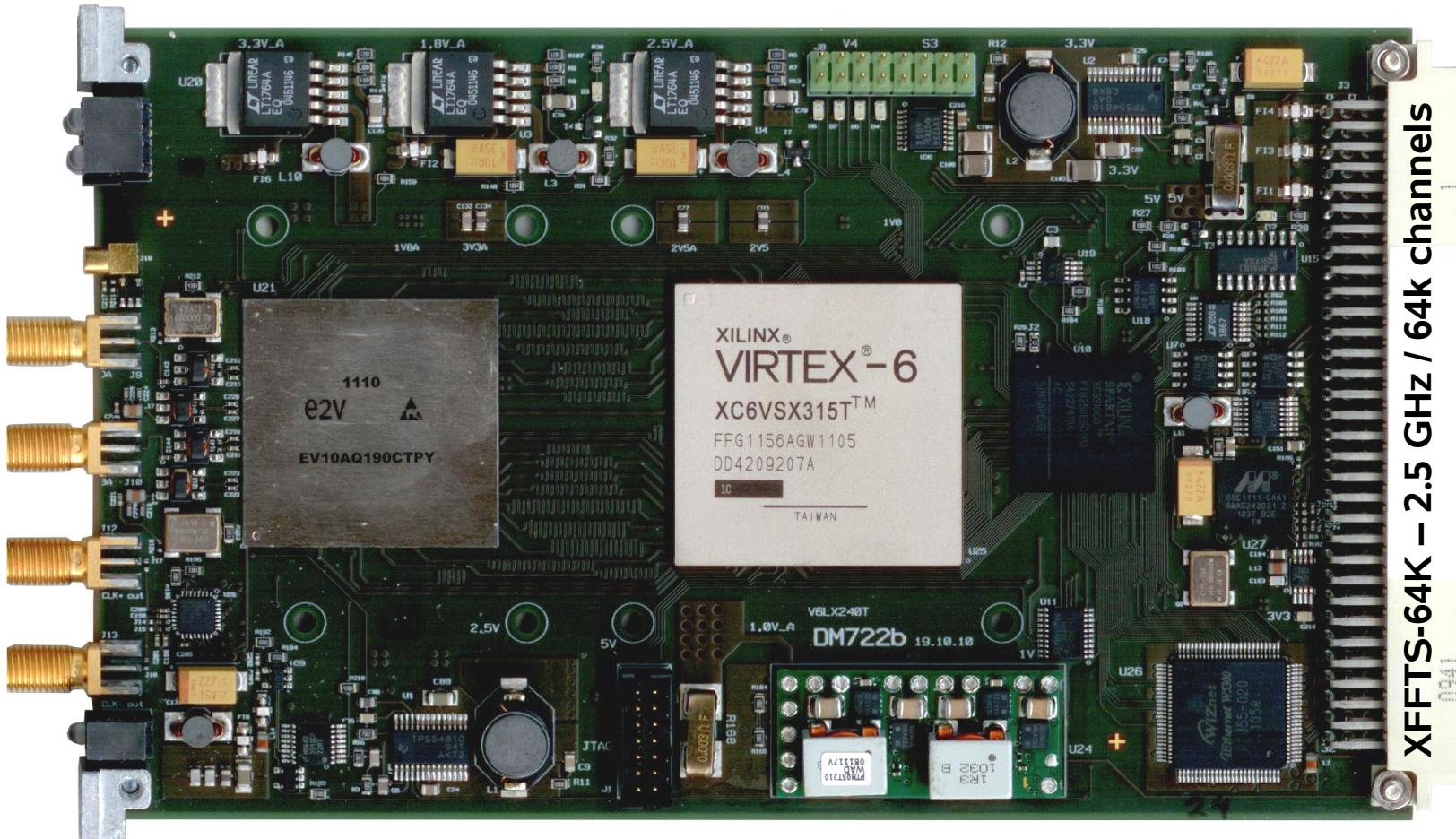


Max-Planck-Institut  
für Radioastronomie

## XFFTS2 :: 64k spectral channels



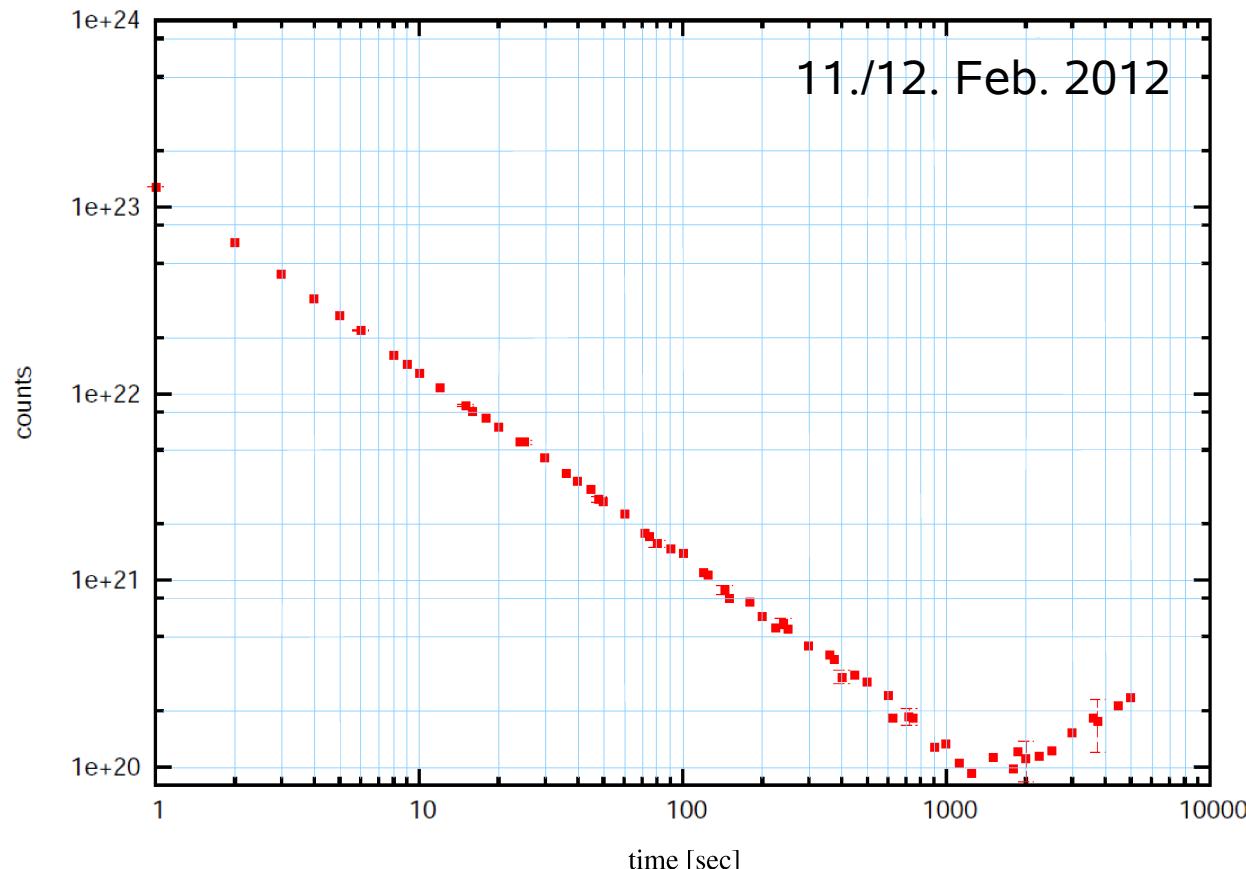
**XFFTS2: 2.5 GHz bandwidth / 65536 channels (ENBW 44 kHz)**





Max-Planck-Institut  
für Radioastronomie

## XFFTS2 :: Allan stability



The spectroscopic Allan variance between two 1.2 MHz broad channels, separated by 1.8 GHz within the band, was determined to be stable on a timescale of >1000 s.



# Fast Fourier Transform Spectrometer (FFTS)

1. Generation: AFFTS – 1.5 GHz BW, 8K channels
2. Generation: XFFTS – 2.5 GHz BW, 32K channels  
XFFTS2 – 2.5 GHz BW, 64K channels
3. Generation: *nextFFTS* – 4 GHz BW, 128K channels  
+ IF sampling (4 – 8 GHz)  
+ digital sideband separation



### 3. Generation : *nextFFTS*

#### Development Goals:

- **4 – 5 GHz instantaneous bandwidth,**
- **128k channels, 4-tap polyphase filter bank,  
ENBW: ~ 35 kHz @ 4 GHz bandwidth**
- **direct IF sampling of the 4 – 8 GHz band**
- **support for digital sideband separation**
- **8-bit ADC, ENOB:  $\geq 6$  bit**
- **100 MBits/s and 1 Gbit/s Ethernet interface**
- **high-resolution continuum mode**
- **power dissipation: < 35 W @ 128k channels**



### 3. Generation : *nextFFTS*



**HMC5448**

**8-BIT, 5 GS/s**  
**ANALOG-TO-DIGITAL CONVERTER**  
► Monolithic ADC

#### Typical Applications

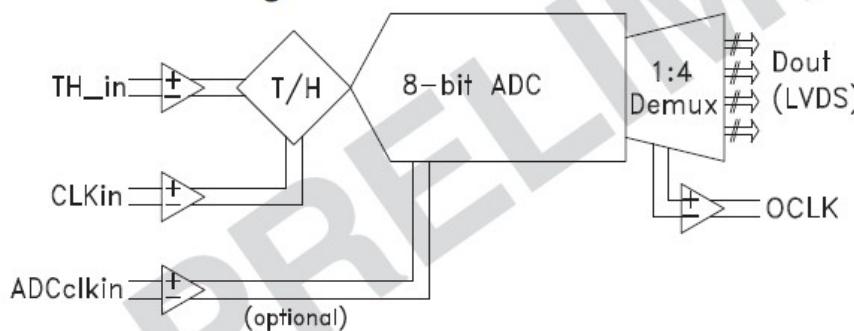
The HMC5448 is ideal for:

- RF Test Instruments and ATE
- Digital Sampling Oscilloscopes
- Radar / Lidar Systems
- Software Defined Radio and Digital Receivers

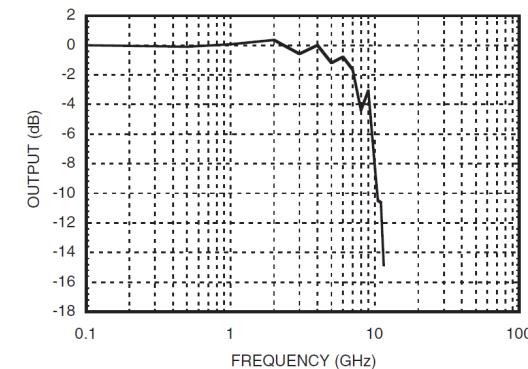
#### Features

- 8-bit Resolution
- >6 GHz Input Bandwidth
- >5 GS/s Sampling Rate
- Direct-coupled Differential Signal, Clock Inputs
- 1:4 LVDS-compatible Output Data Demux
- Synchronous Output Clock, Over-range Bit
- 16 x 16 pin (17 x 17 mm<sup>2</sup>) BGA Package

#### Functional Diagram



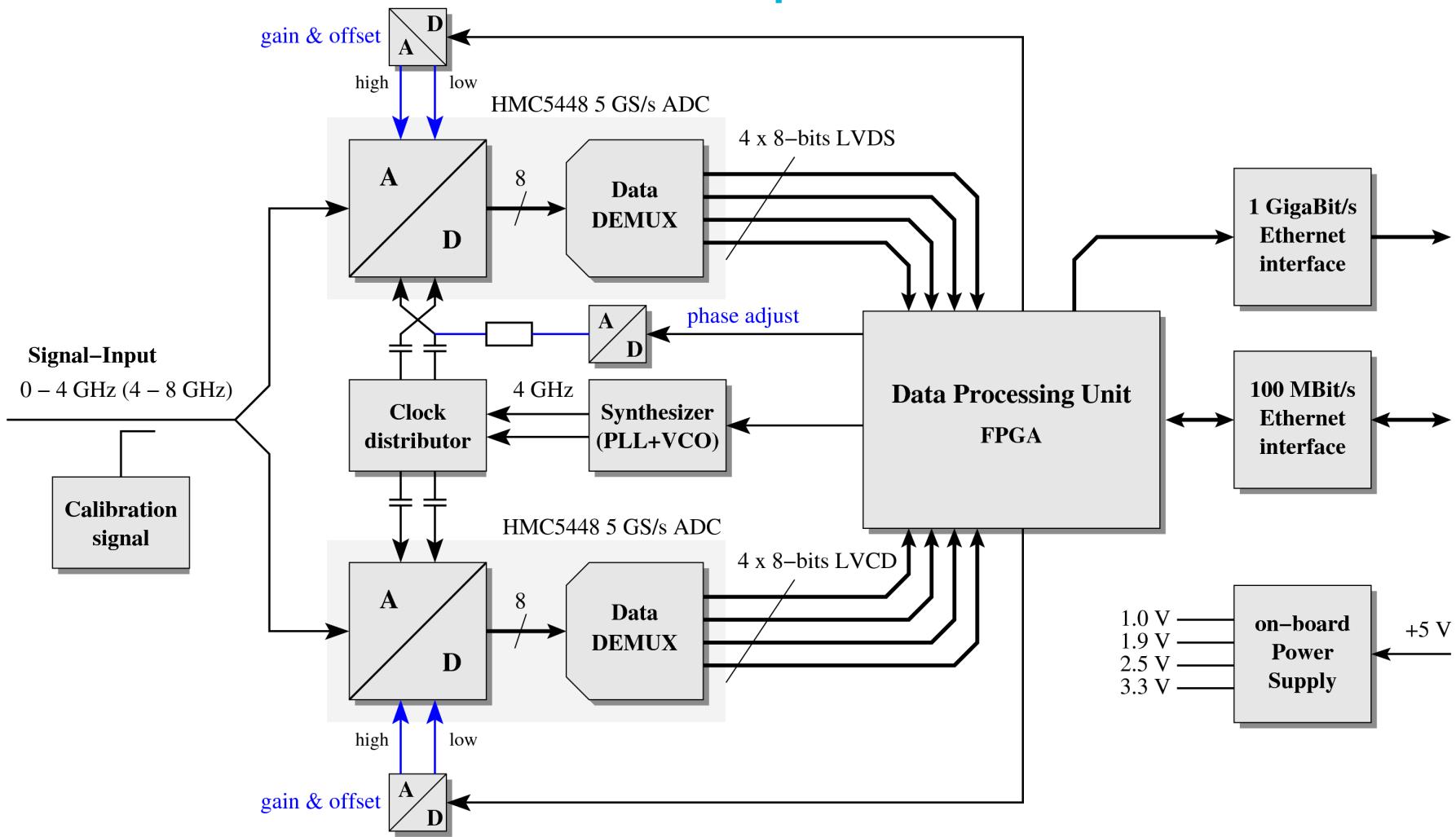
#### Input Bandwidth [1]





### 3. Generation : *nextFFTS*

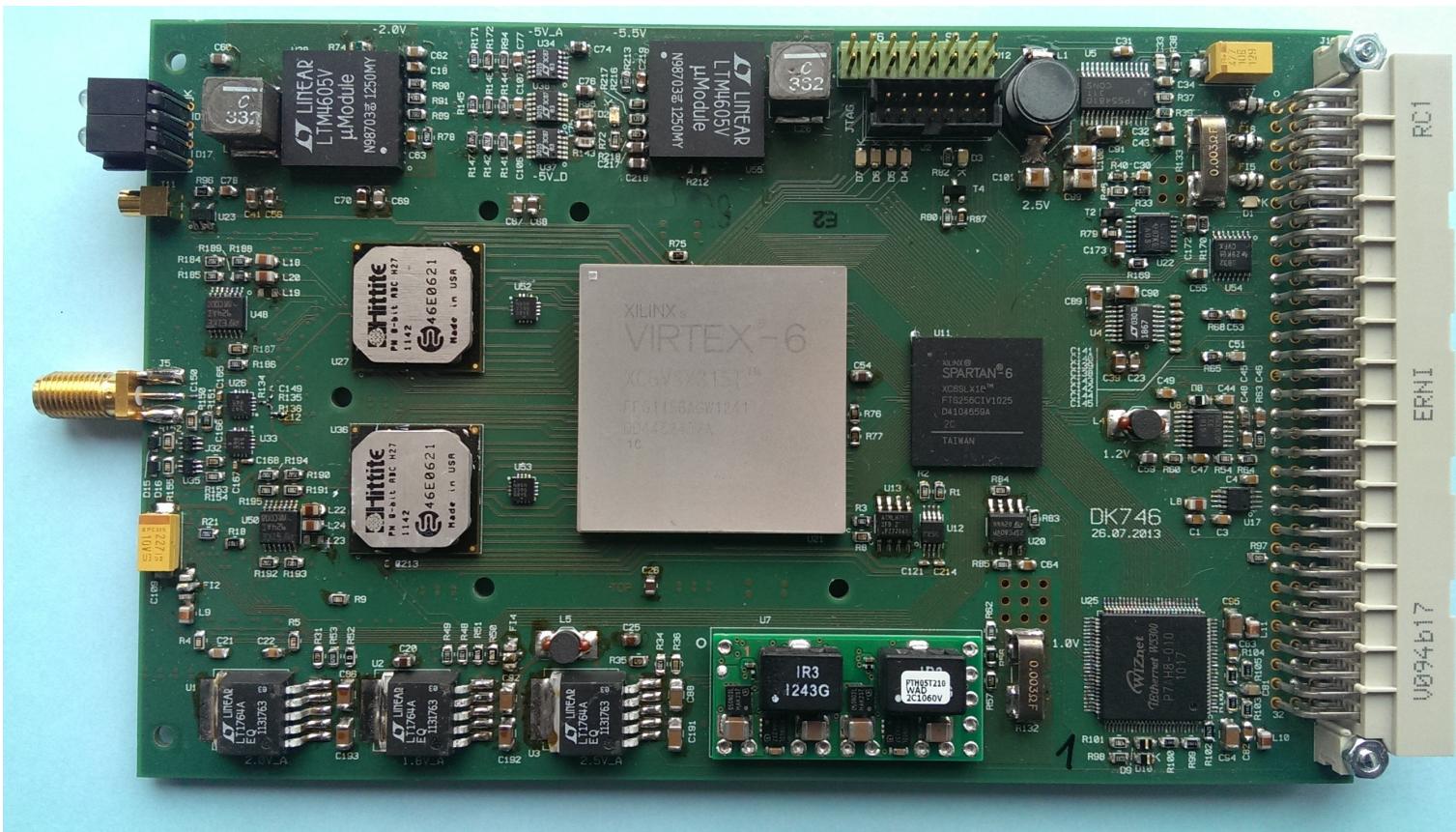
#### Concept





### 3. Generation : *nextFFTS*

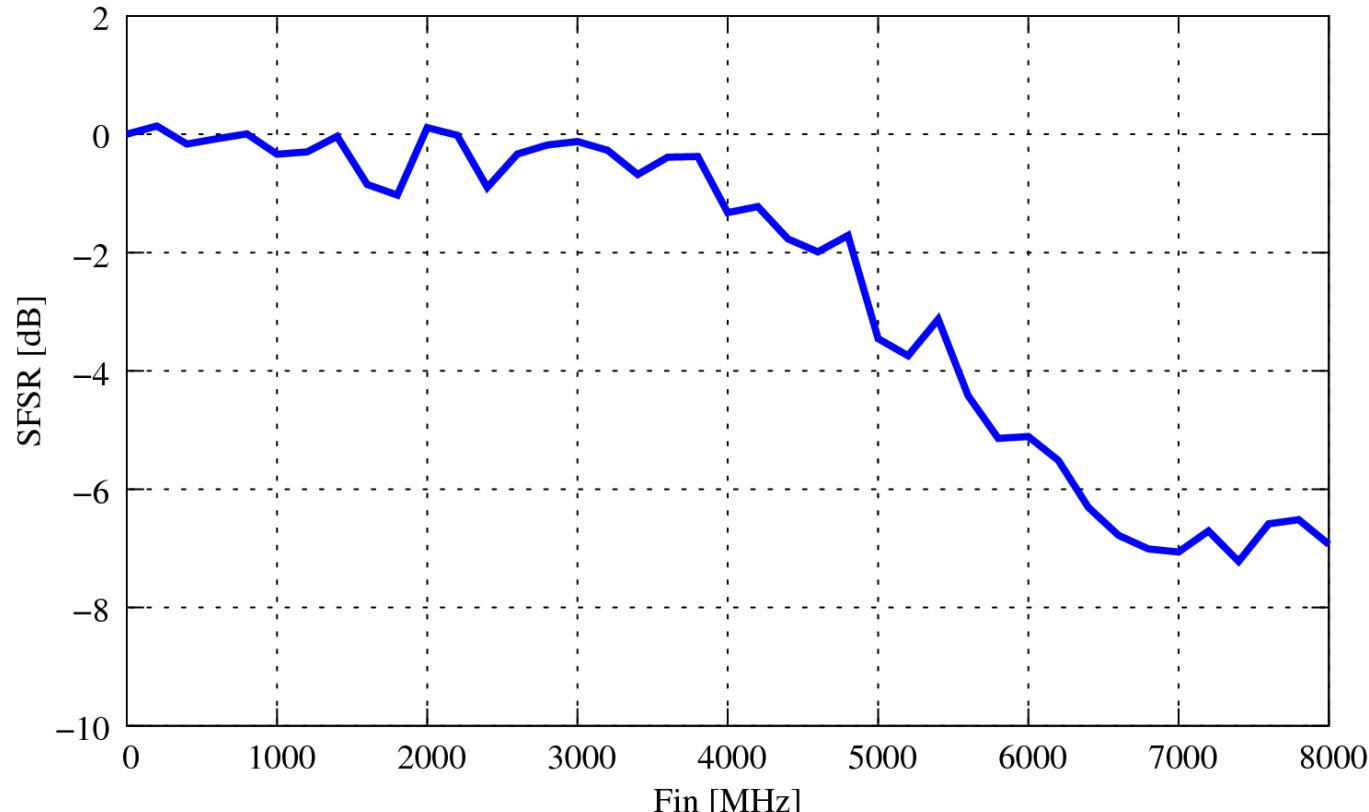
## Photo of the Prototype Board (December 2013)





### 3. Generation : *nextFFTS*

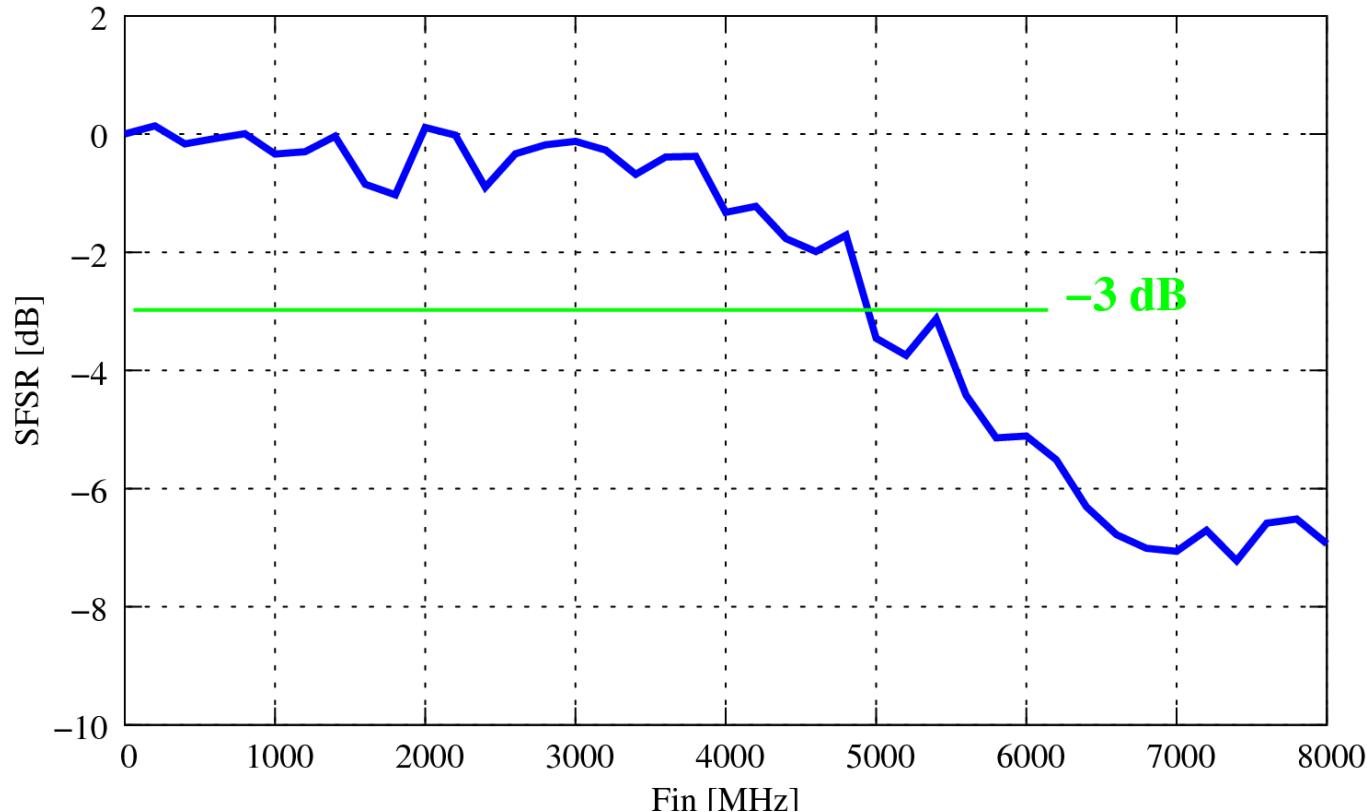
Measured Normalized Full Power Input Bandwidth (fsamp = 4 GHz)





### 3. Generation : *nextFFTS*

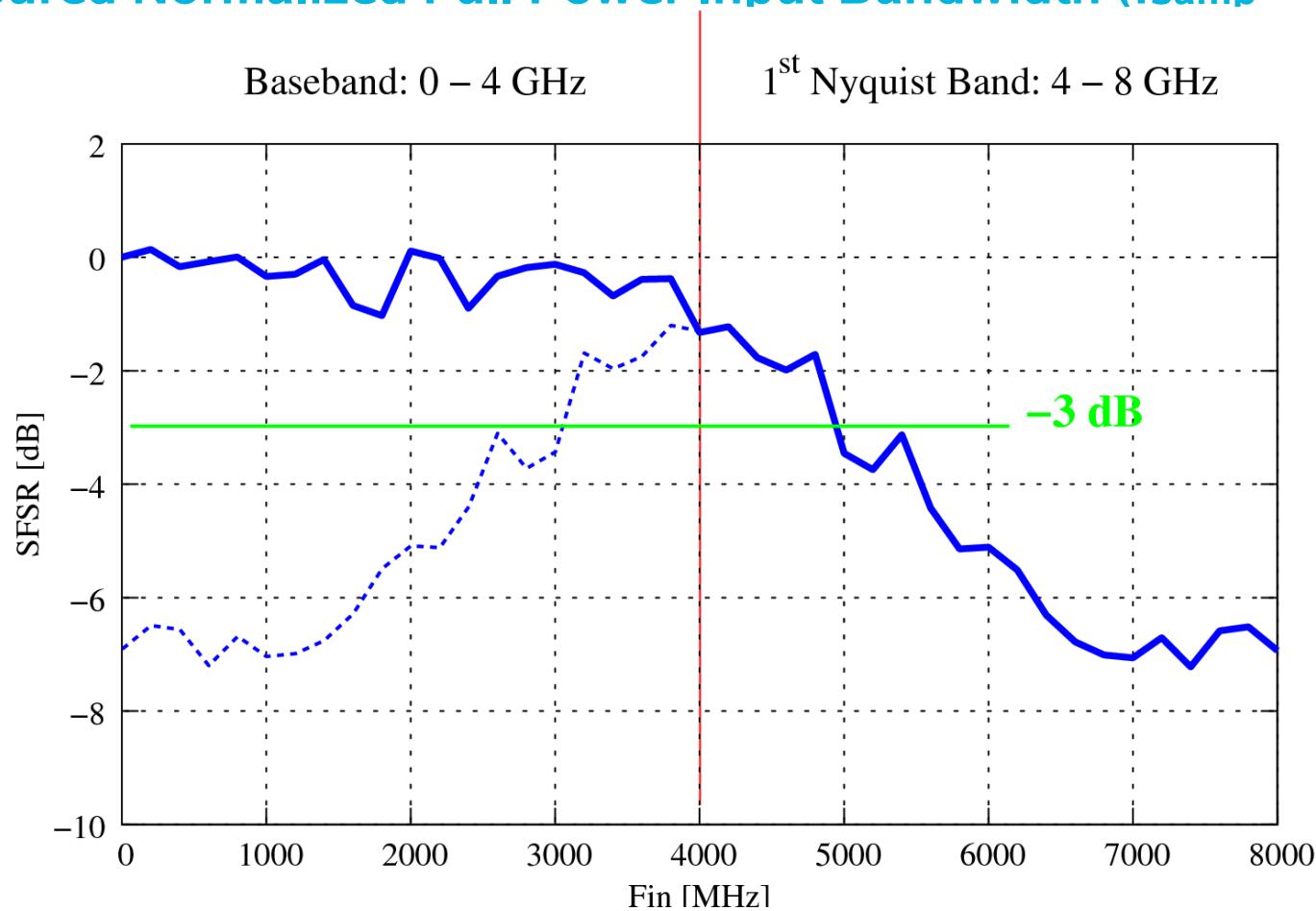
Measured Normalized Full Power Input Bandwidth (fsamp = 4 GHz)





### 3. Generation : *nextFFTS*

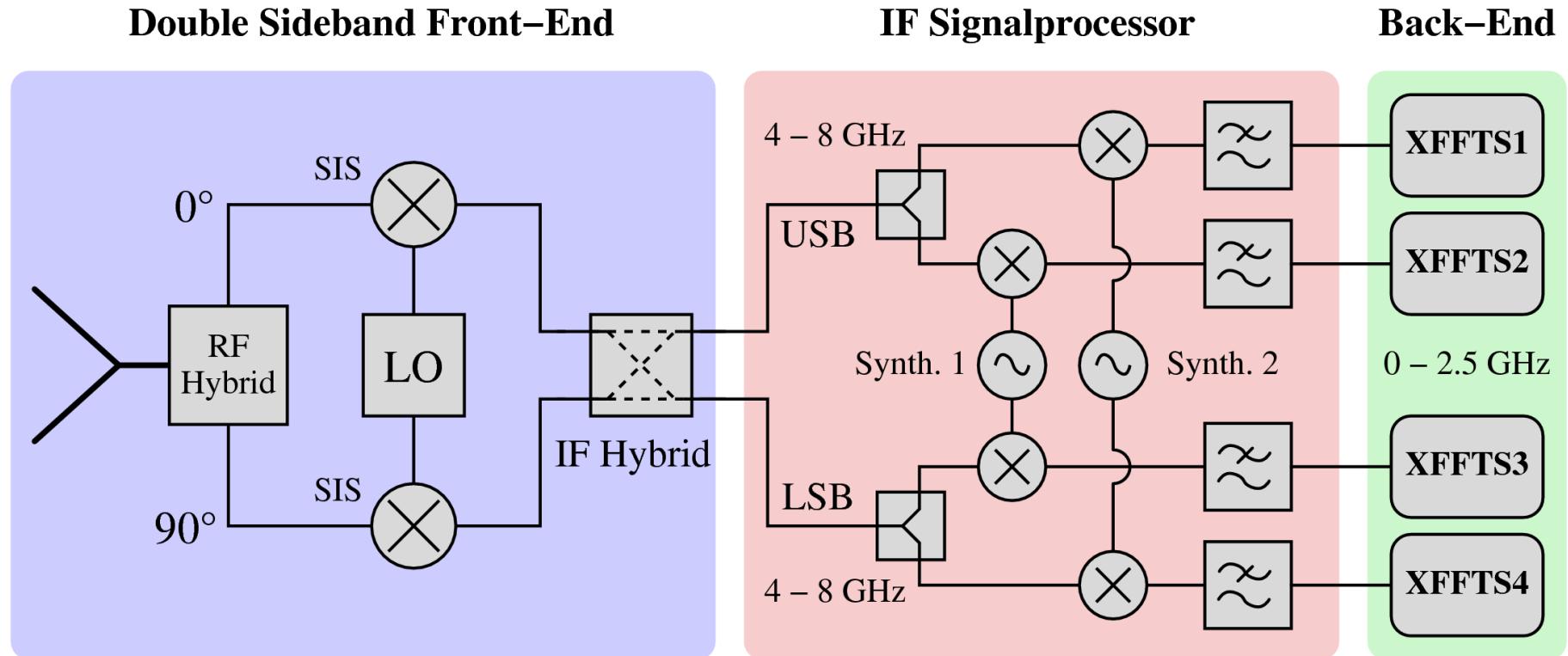
#### Measured Normalized Full Power Input Bandwidth (fsamp = 4 GHz)





### 3. Generation : *nextFFTS*

#### Classical Heterodyne system :: APEX / FLASH<sup>+</sup>

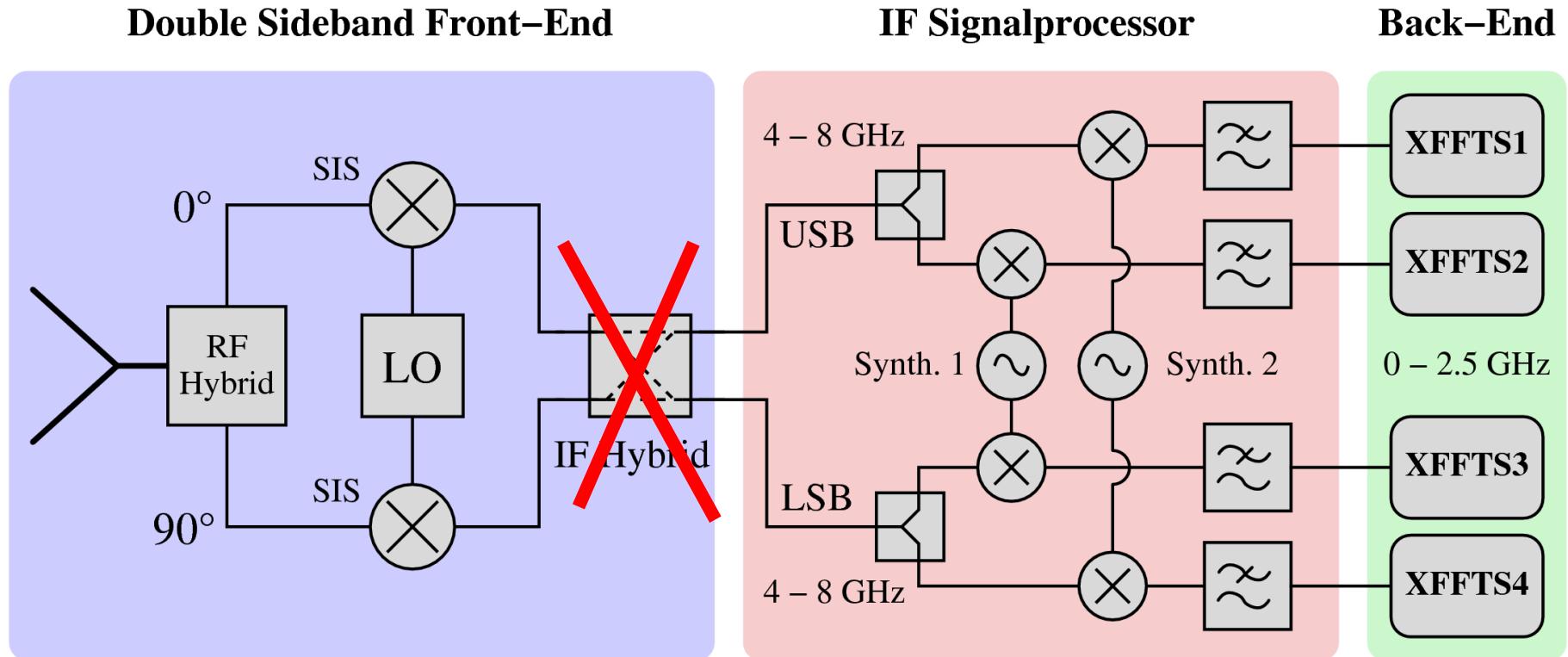


- Problem: amplitude and phase imbalances limit the sideband separation ratio to 10 - 15 dB.
- An analog IF system can become very complex, expensive and unstable.



### 3. Generation : *nextFFTS*

#### Classical Heterodyne system :: APEX / FLASH<sup>+</sup>

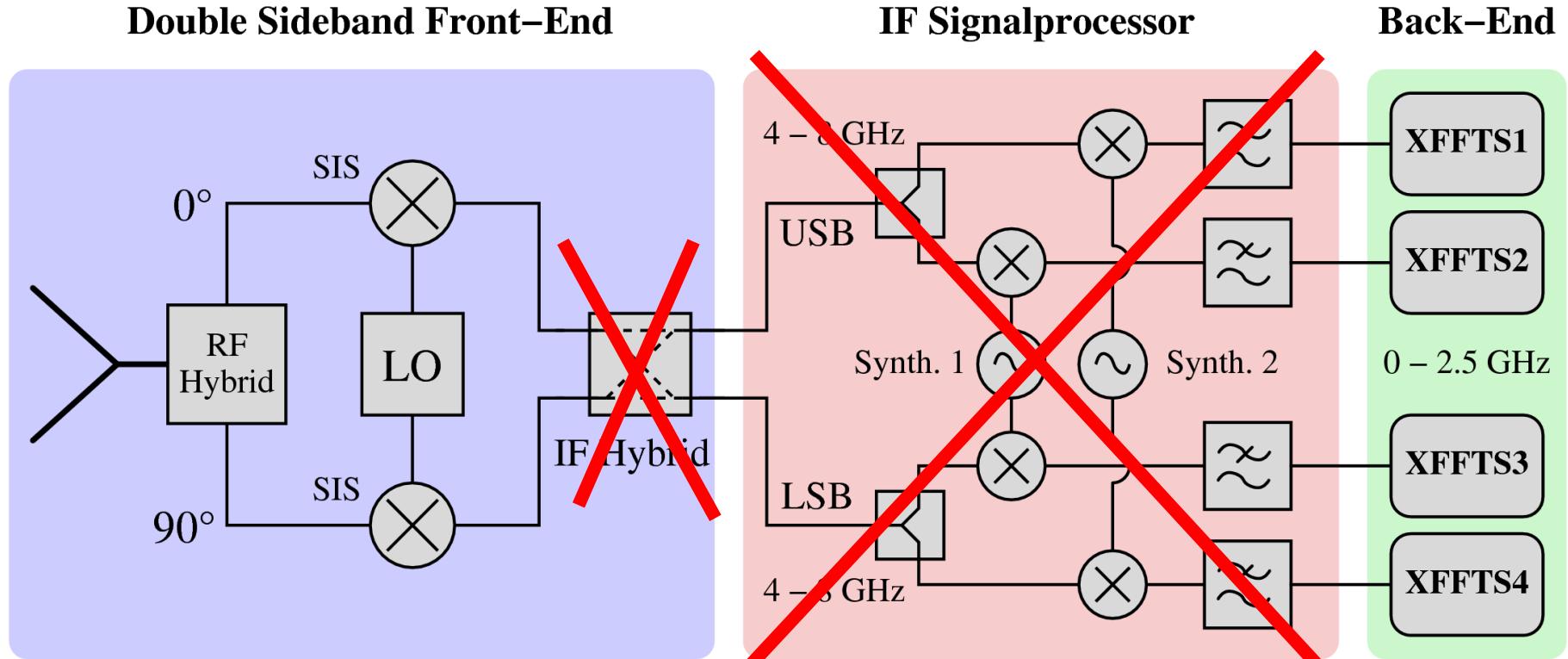


- Problem: amplitude and phase imbalances limit the sideband separation ratio to 10 - 15 dB.
- An analog IF system can become very complex, expensive and unstable.



### 3. Generation : *nextFFTS*

#### Classical Heterodyne system :: APEX / FLASH<sup>+</sup>



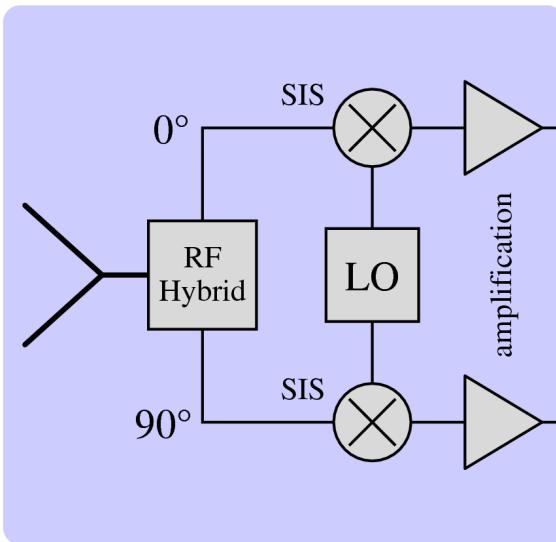
- Problem: amplitude and phase imbalances limit the sideband separation ratio to 10 - 15 dB.
- An analog IF system can become very complex, expensive and unstable.



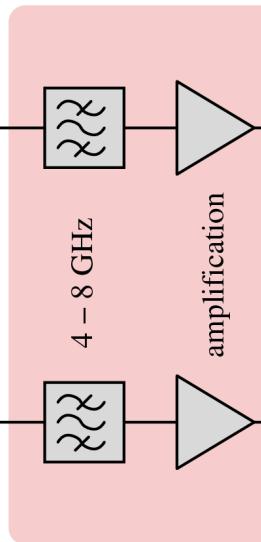
### 3. Generation : *nextFFTS*

## New concept for Heterodyne Receiver

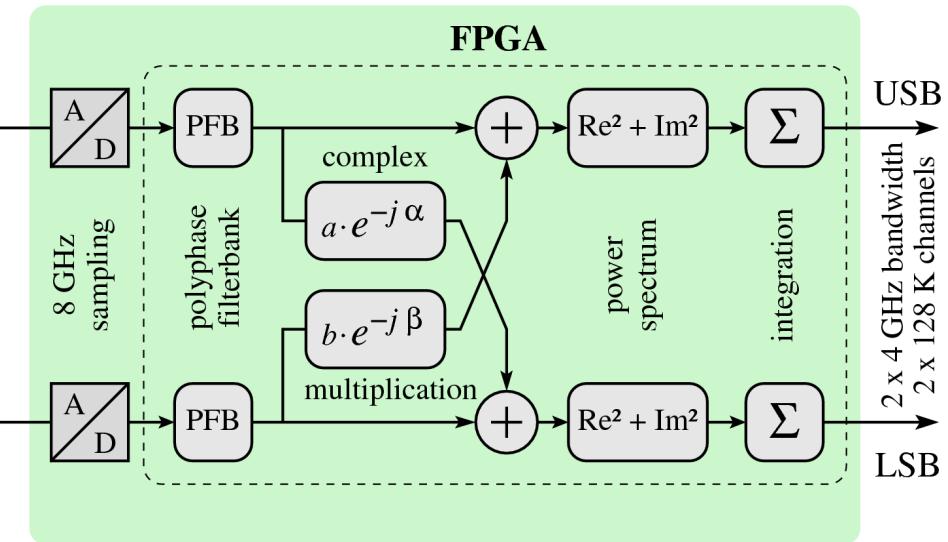
Double Sideband Front-End



IF Processor



Back-End: Sideband Separating FFTS



- Front-End without an IF-hybrid.
- 90° phase shift is implemented by the FFTS for each spectral channel.
- IF processor as simple as possible.
- Undersampling techniques (digital down mixing) instead of analog base-band mixing.
- IF sampling at 8 GS/s to process an instantaneous bandwidth of 4 GHz.



## FFT-Spectrometer :: Summary

### Advantages of our new generation of compact FFT spectrometers:

- ✓ FFTS offer high instantaneous bandwidth up to > 4 GHz with many thousands frequency channels, thus offering wide-band observations with high spectral resolution without the complexity of the IF processing in a hybrid configuration.
- ✓ They provide very high stability by exclusive digital signal processing. Allan stability times of > 1000 seconds have been demonstrated routinely.
- ✓ Our optimized polyphase FFT signal processing pipeline provides a nearly loss-free time to frequency transformation with significant reduced frequency scallop, less noise bandwidth expansion, and faster side lobe fall-off.
- ✓ Field-operations of our FFTS over the last 8 years have proven to be very reliable, with calibration- and aging-free digital processing boards.
- ✓ Low space and power requirements – thus safe to use at high altitude (e.g. APEX at 5100-m) as well as on spacecrafts (SOFIA) and future satellites (Millimetron?).
- ✓ Production cost are low compared to traditional spectrometers through use of only commercial components.



Max-Planck-Institut  
für Radioastronomie

## FFTS/XFFTS :: Contact, Distribution

DIGITALLABOR  
0100110 01000110 01010100 01010011 – 01000010 01001011

### Contact:

For further information about the MPIfR FFT spectrometer,  
future developments and applications, please contact

Dr. Rolf Güsten ([ruesten@mpifr.de](mailto:ruesten@mpifr.de)) or

Prof. Bernd Klein ([bklein@mpifr.de](mailto:bklein@mpifr.de)) at the

Max-Planck-Institut für Radioastronomie in Bonn, Germany.



### Distribution:



<http://www.radiometer-physics.de>